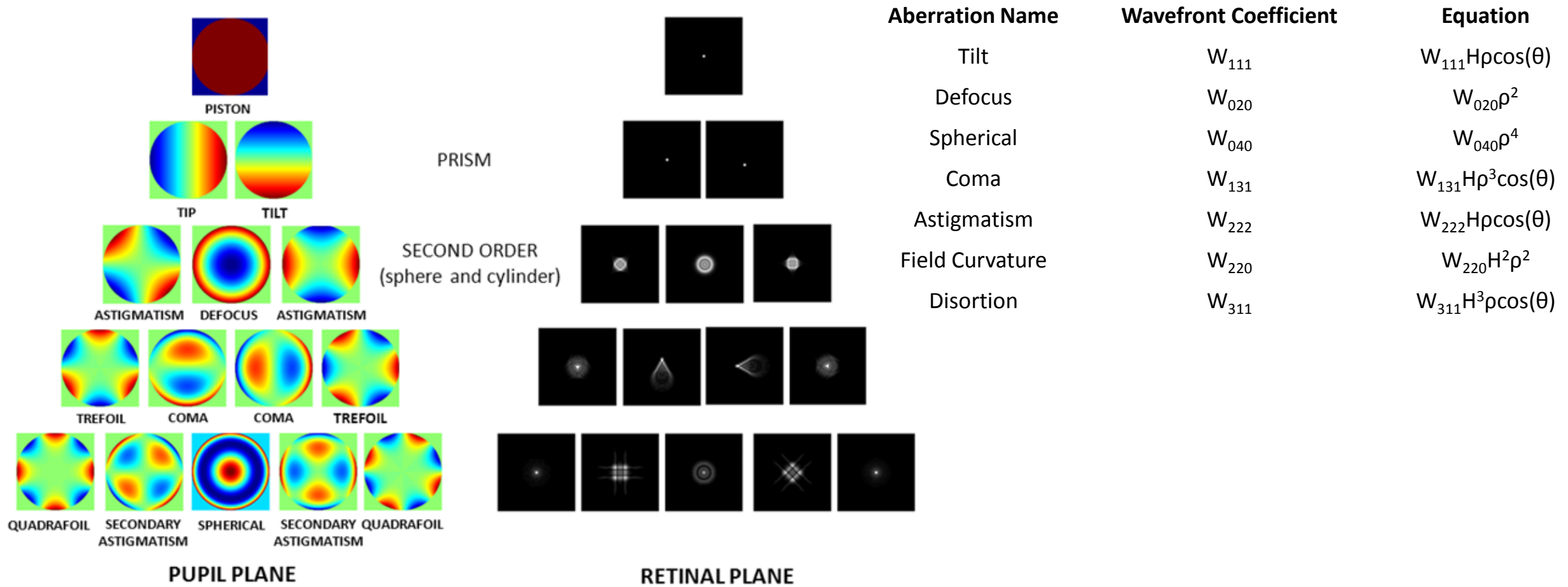


# Aberrations and light sources

- Last class
  - Diffraction and imaging
  - Numerical aperture, diffraction limited spot, resolution limit
  - Aberrations
- This class
  - Finish aberrations
  - Light sources
  - Phase contrast imaging

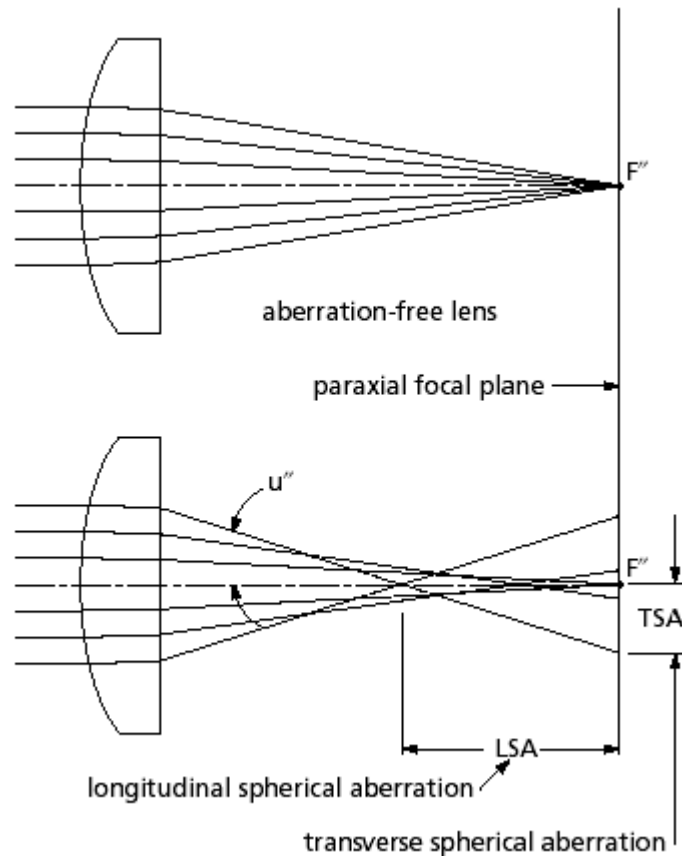
# Aberrations are well mapped



A

B

# Aberrations are imperfections in collecting rays



## Zernike polynomials

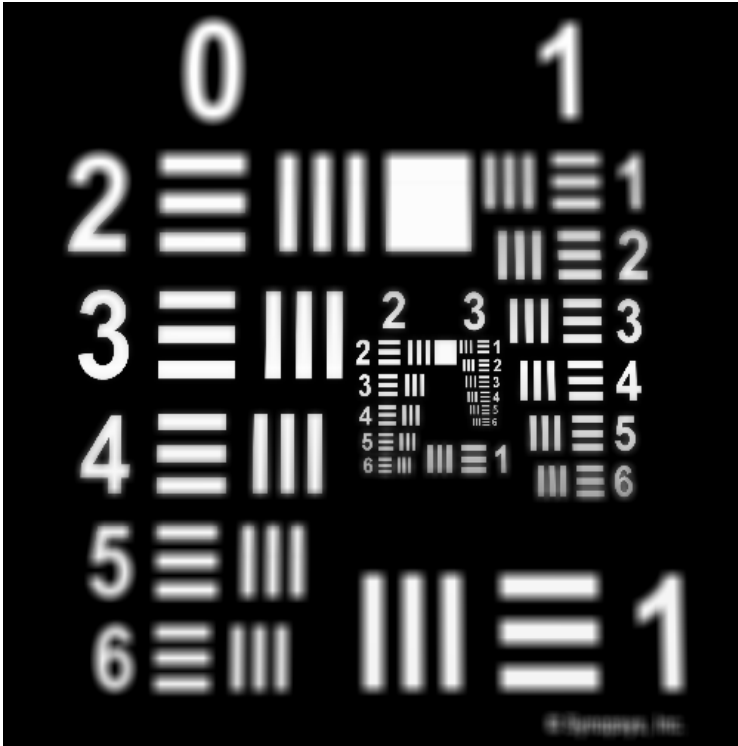
$$Z_n^m(\rho, \varphi) = R_n^m(\rho) \cos(m\varphi)$$

$$Z_n^{-m}(\rho, \varphi) = R_n^m(\rho) \sin(m\varphi),$$

$$R_n^m(\rho) = \sum_{k=0}^{\frac{n-m}{2}} \frac{(-1)^k (n-k)!}{k! \left(\frac{n+m}{2} - k\right)! \left(\frac{n-m}{2} - k\right)!} \rho^{n-2k}$$

Zernike polynomials have a complicated form, but don't worry about those equations

# Tilt aberrations

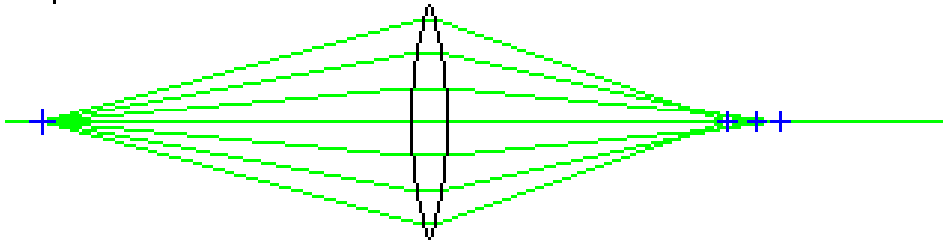


First order aberration. Usually means your sample is not placed above the objective well.

Completely correctable.

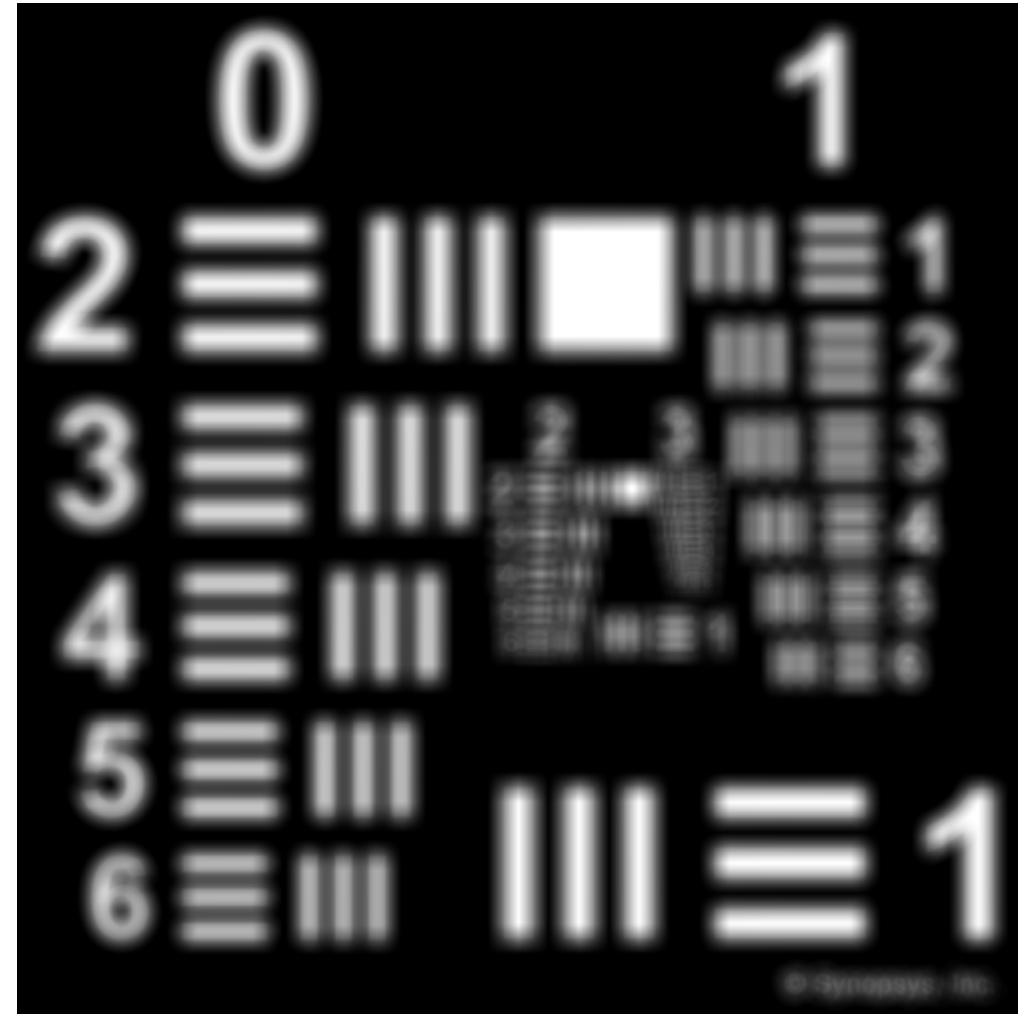
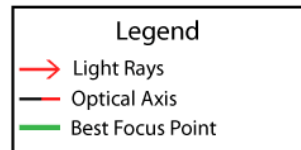
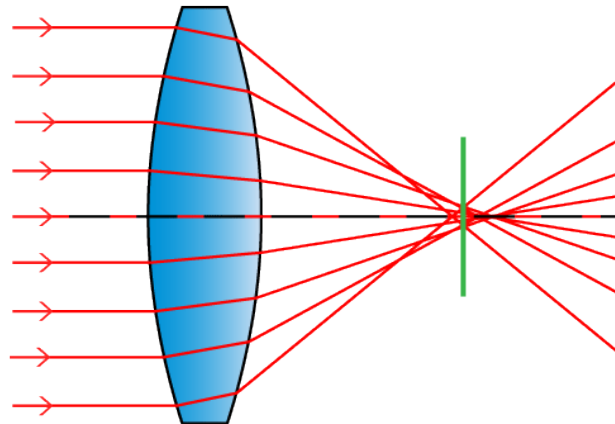
# Spherical aberrations

Spherical Aberration



Arise from the spherical curvature of the lens

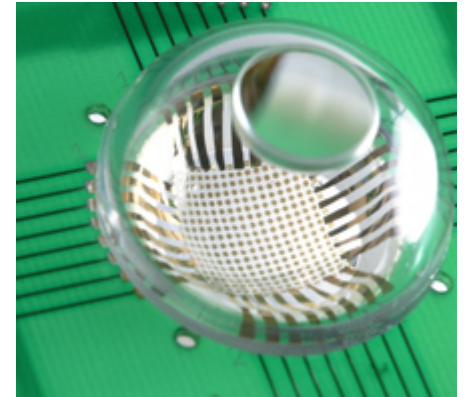
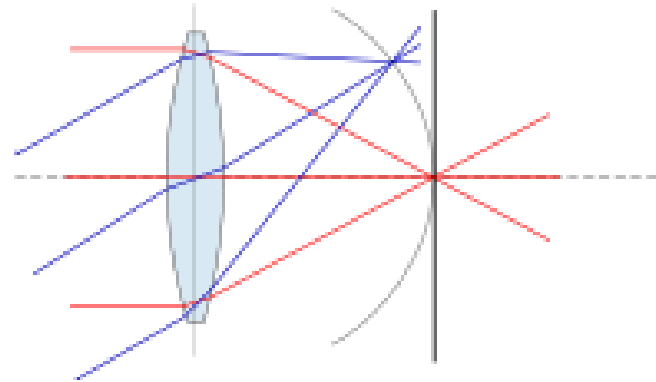
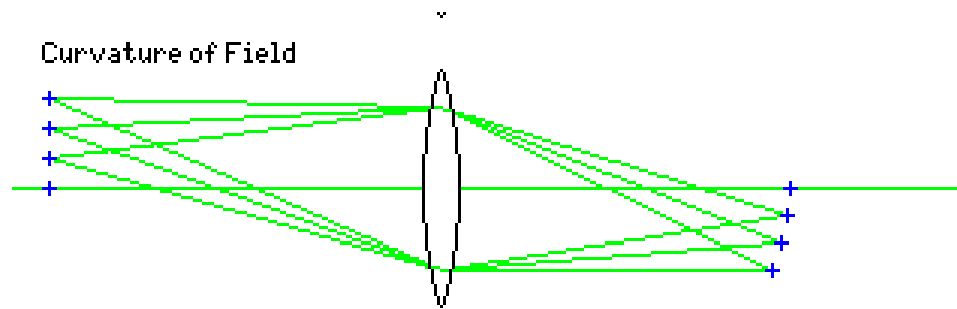
Lens with Spherical Aberration



Corrections:

- Use multiple lenses with different indices
- Use aspheric lenses

# Curvature of field



Arise because we focus onto a flat surface.  
The lens will ideally focus onto a spherical detector.

Cameras are hard to make as spheres.

Plan lenses correct for flat plane imaging

Limits overall field number (current limit 20-25 mm)

Objective Correction for Field Curvature

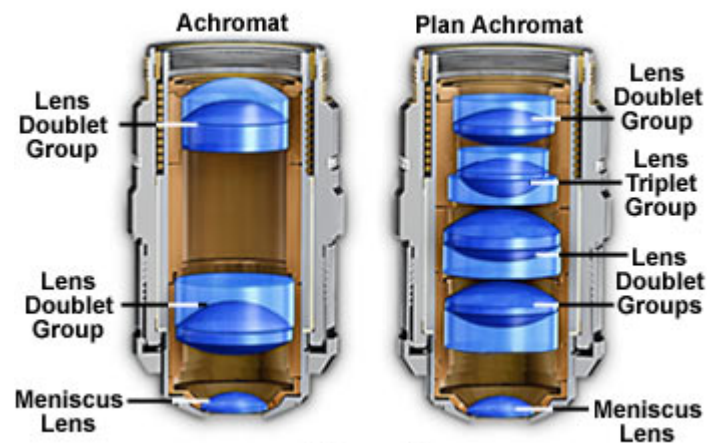


Figure 2

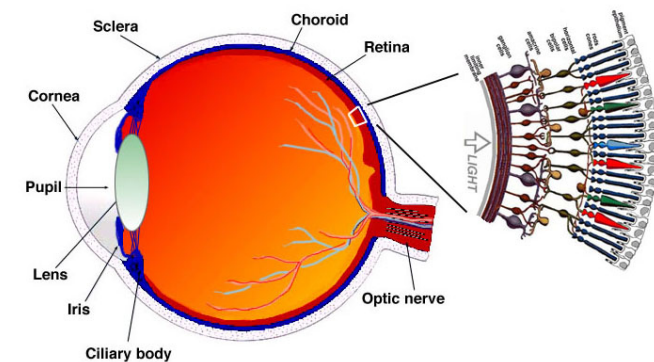
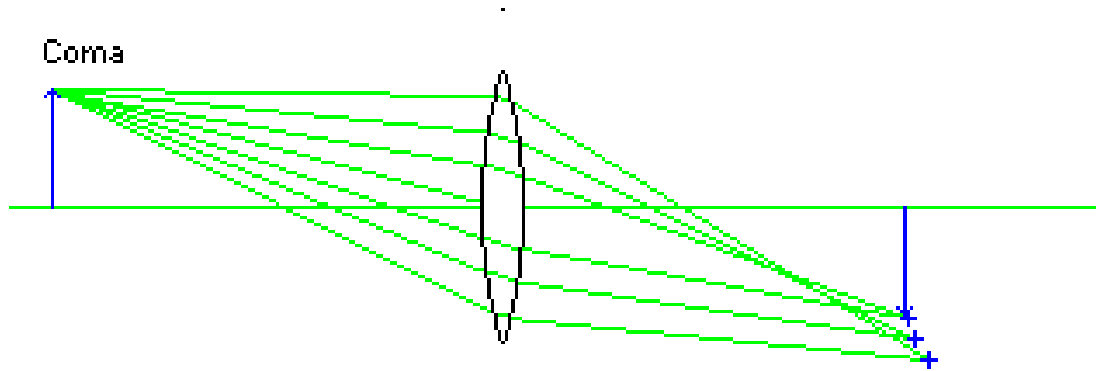


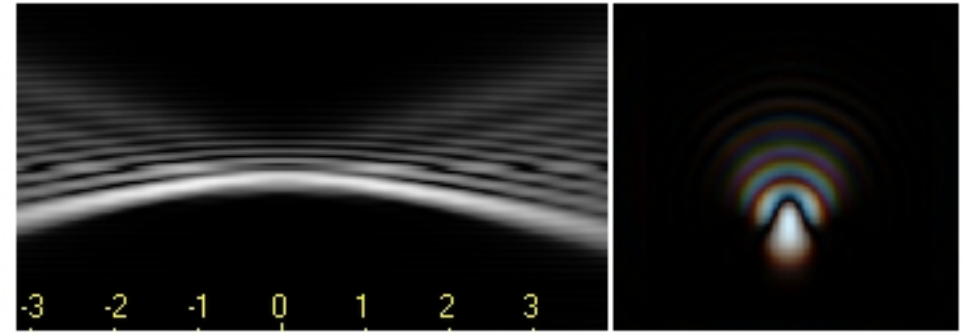
Fig. 1.1. A drawing of a section through the human eye with a schematic enlargement of the retina.

# Coma



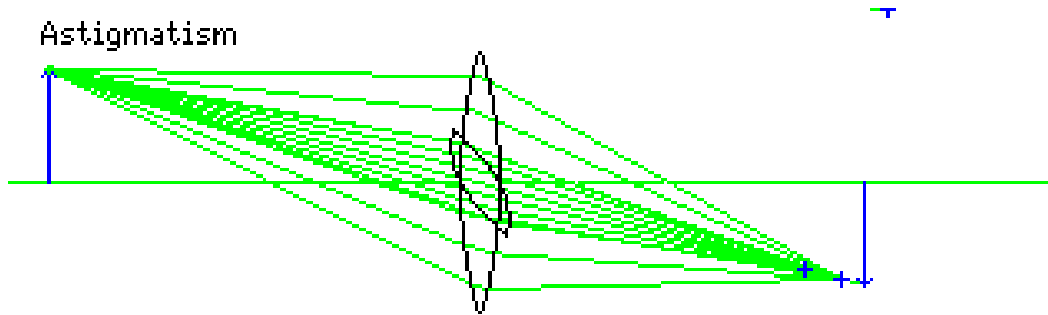
Dependent on lens shape.  
Easily seen when a lens is tilted relative to the incoming light

Usually means that a lens is out of alignment





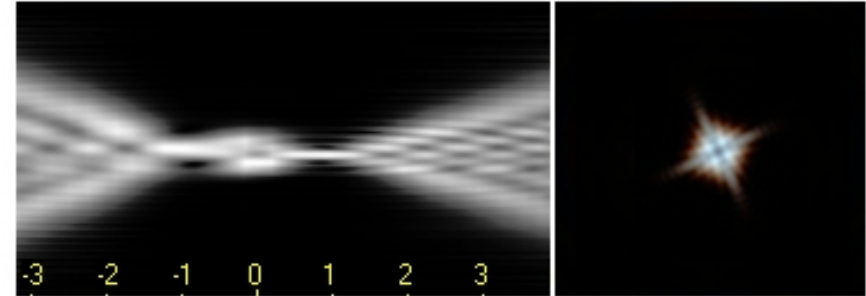
# Astigmatism



Proportional to the diameter of optics

Changes from vertical to horizontal going in and out of focus

Original	Compromise
aio	aio
Horizontal Focus	Vertical Focus
aio	aio



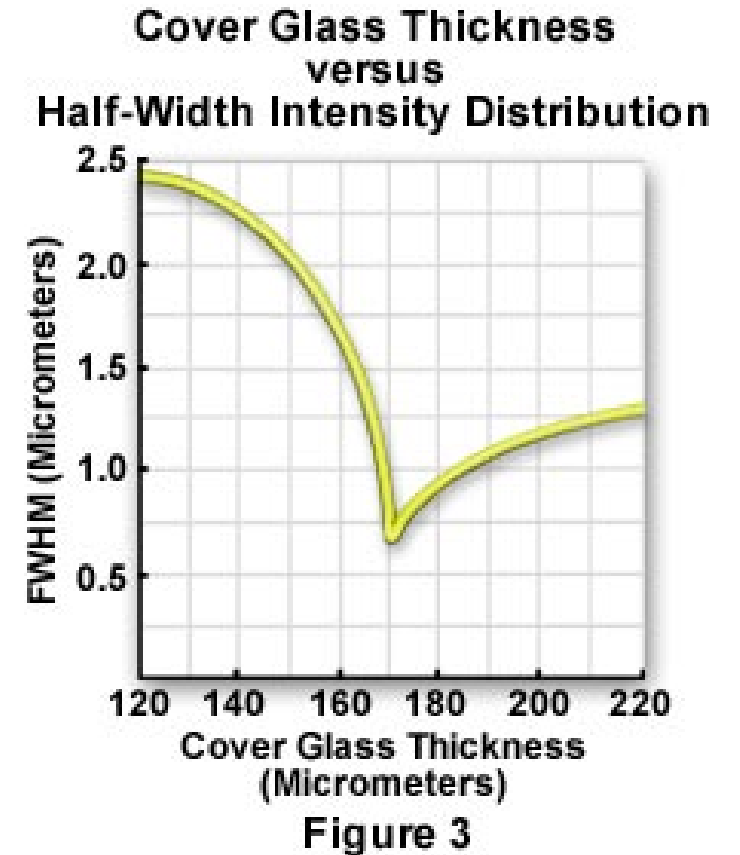
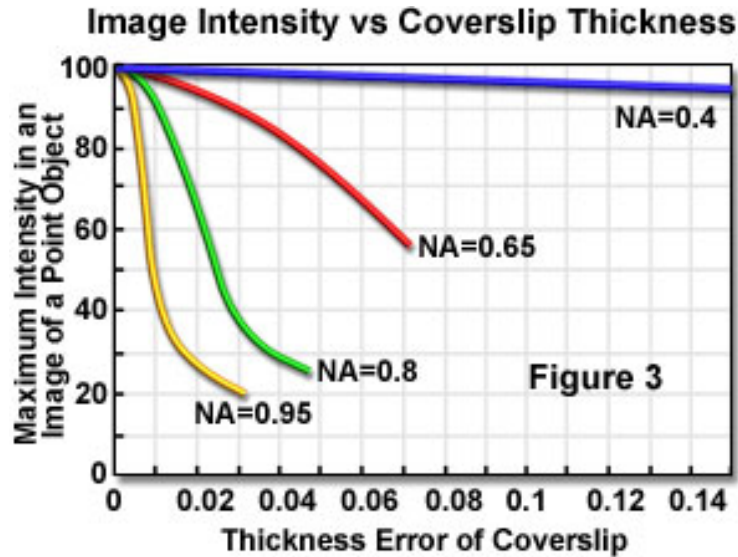
Another signature of mis-aligned optics

# New objectives, coverslips

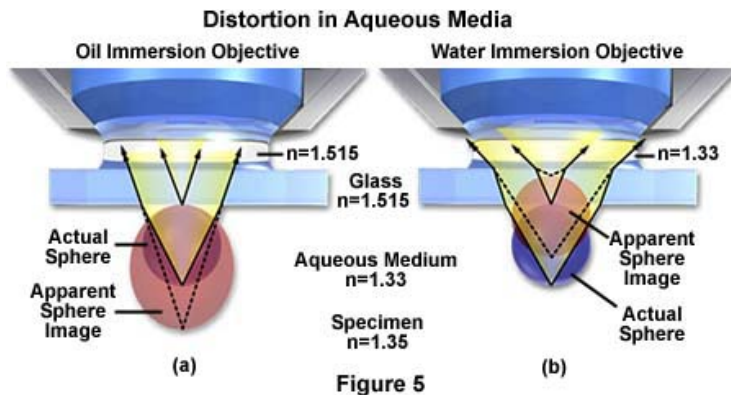
The transition from sample to coverslip to air/water to glass can cause serious distortions, especially at high NA

Most high NA objectives are made with correction collars that will adjust for this issue

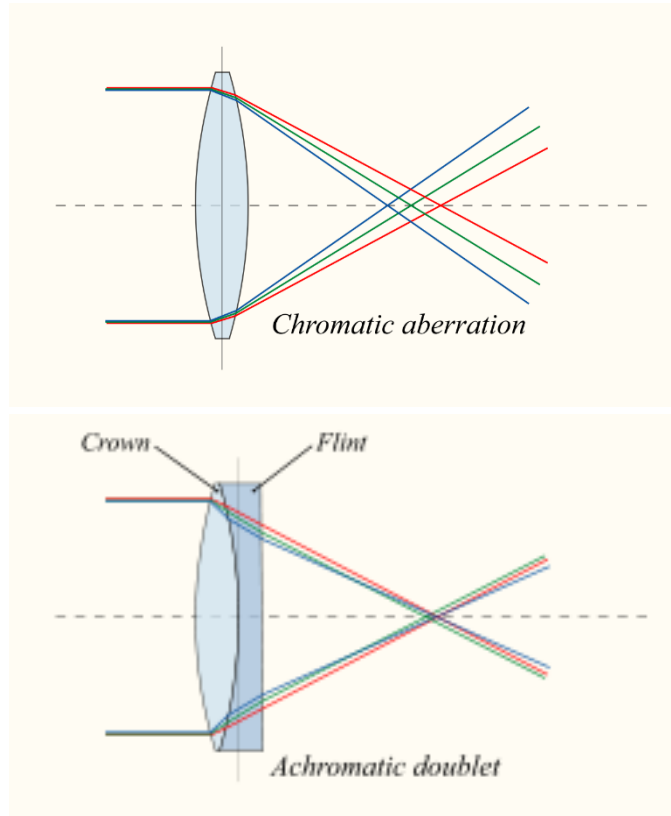
Make sure objective is coverslip corrected for inverted imaging!



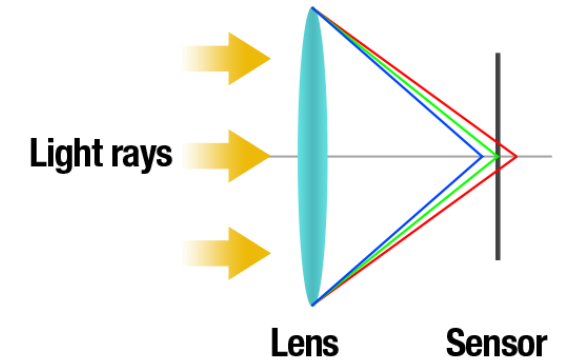
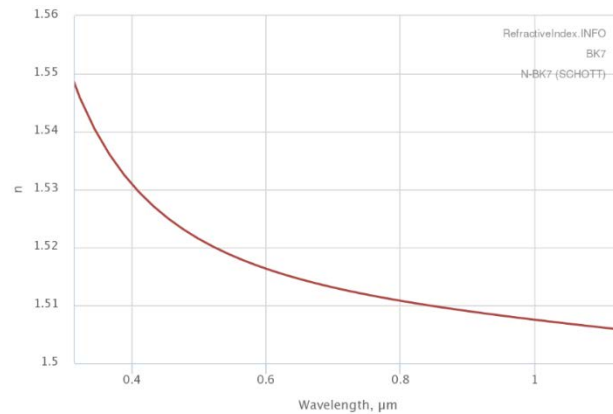
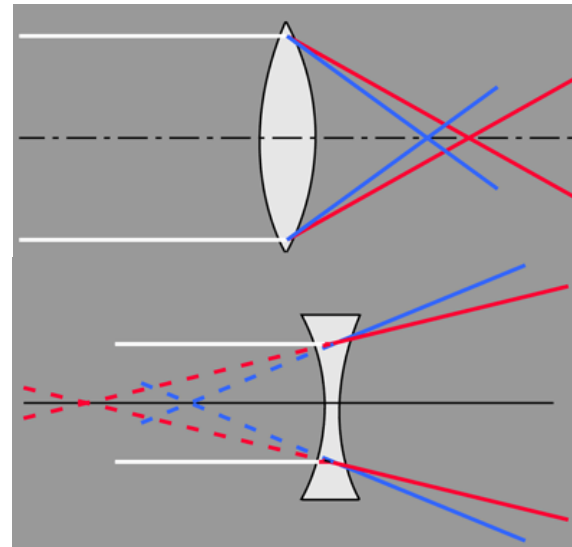
- No. 0 – 0.085 to 0.13 mm thick
- No. 1 – 0.13 to 0.16 mm thick
- No. 1.5 – 0.16 to 0.19 mm thick
- No. 1.5H – 0.17 to 0.18 mm thick
- No. 2 – 0.19 to 0.23 mm thick
- No. 3 – 0.25 to 0.35 mm thick
- No. 4 – 0.43 to 0.64 mm thick



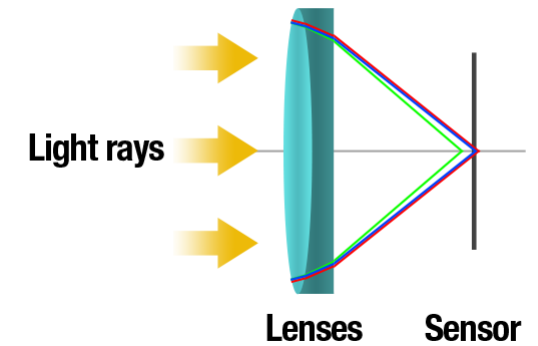
# Chromatic aberrations



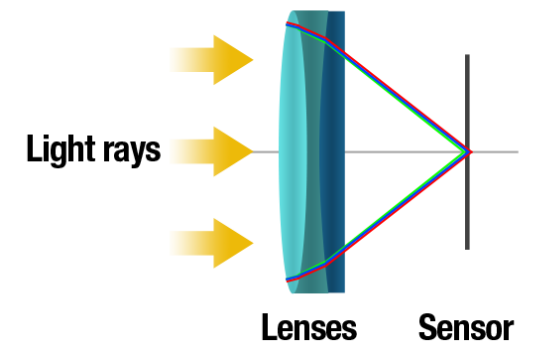
Use multiple lenses of varying indices.



**Achromatic lens design**



**Apochromatic lens design**

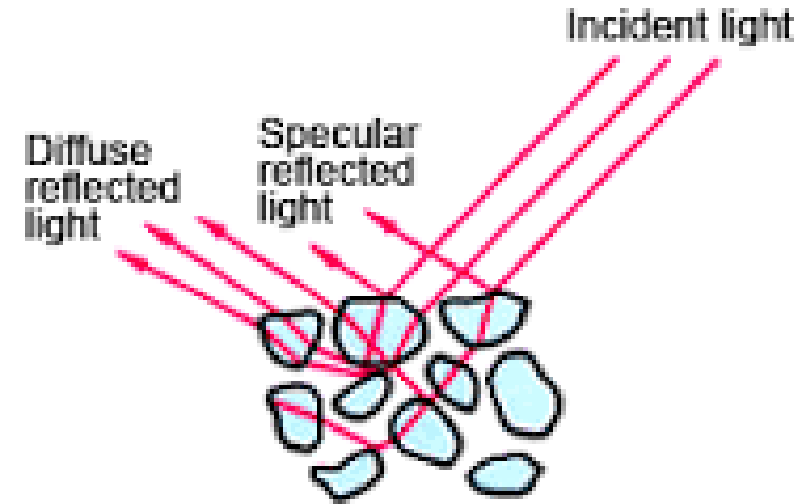


$$\frac{1}{f} = (n - 1) \left[ \frac{1}{R_1} - \frac{1}{R_2} + \frac{(n - 1)d}{nR_1R_2} \right],$$

$$\frac{1}{f_{tot}} = \frac{1}{f_1} + \frac{1}{f_2}$$

# Scattering

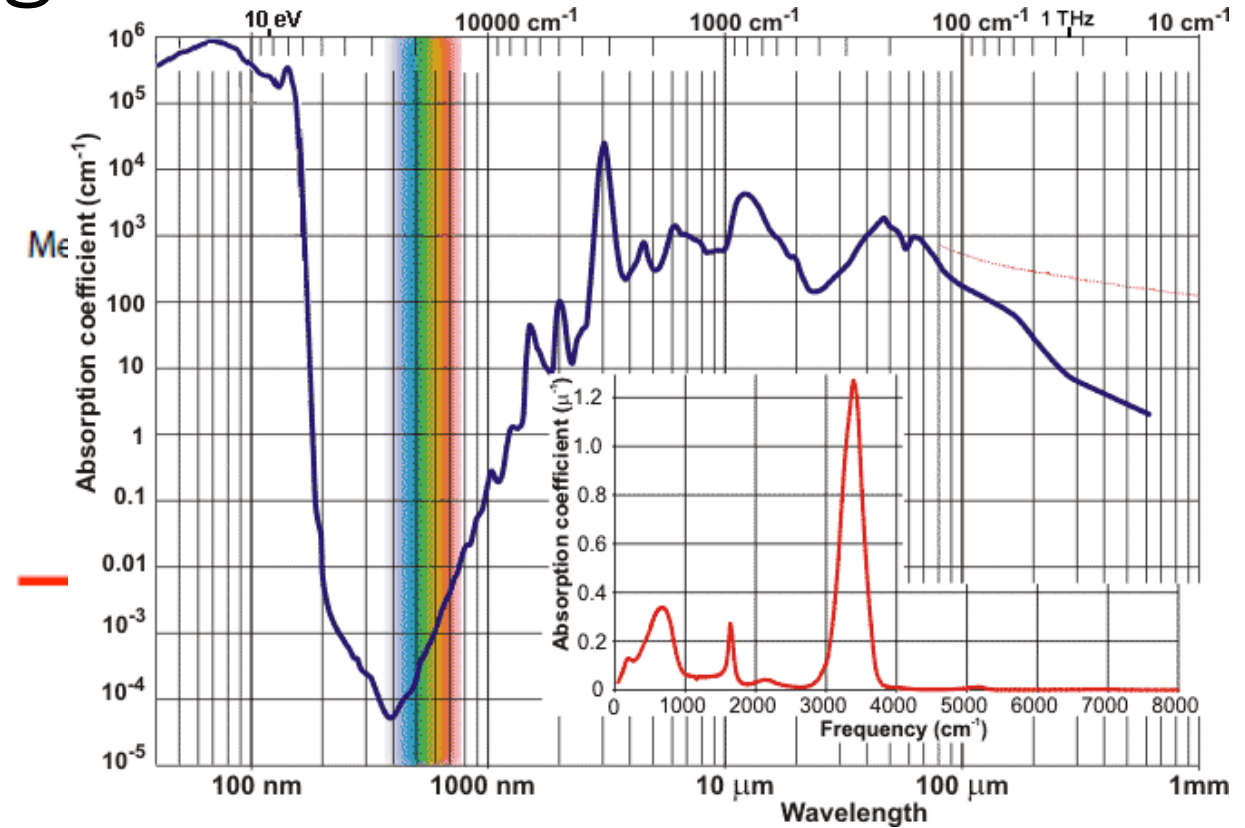
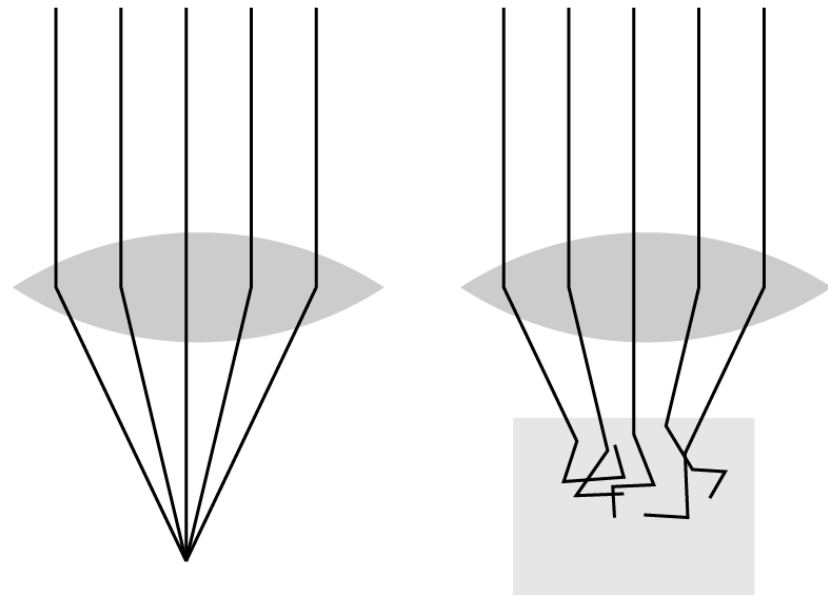
- Scattering is reflection, but off very small particles
- Wavelength in = wavelength out
- Light is redirected with angle probability  $\cos^2\theta$
- Scattering in biology is mostly due to lipid bilayers ( $n \sim 1.5$ )



$$\sigma_s = \frac{2\pi^5}{3} \frac{d^6}{\lambda^4} \left( \frac{n^2 - 1}{n^2 + 2} \right)^2$$



# Scattering changes the directions of light, makes imaging tough

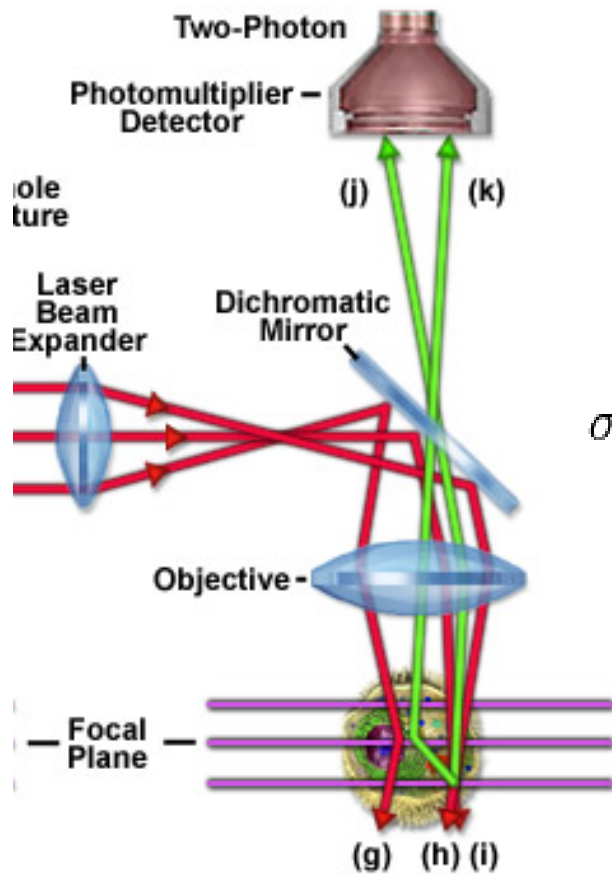


Mean free path  $\sim 100\mu\text{m}$  in biological tissue

$$\sigma_s = \frac{2\pi^5}{3} \frac{d^6}{\lambda^4} \left( \frac{n^2 - 1}{n^2 + 2} \right)^2$$

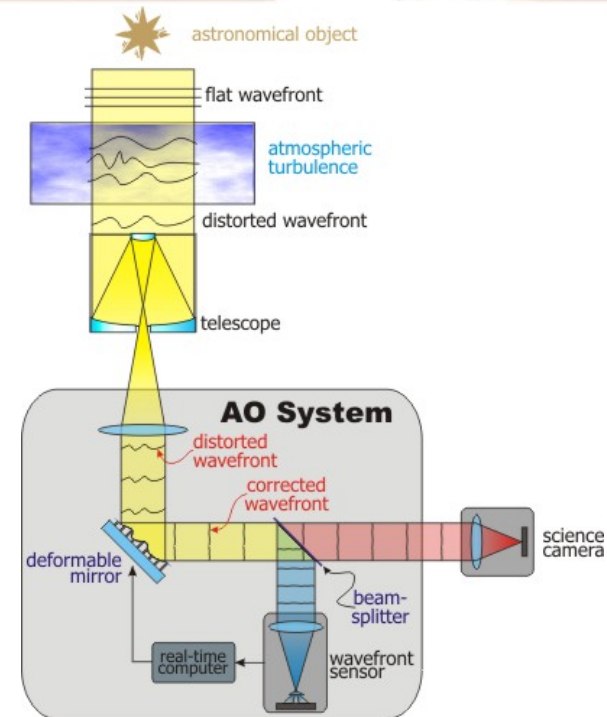
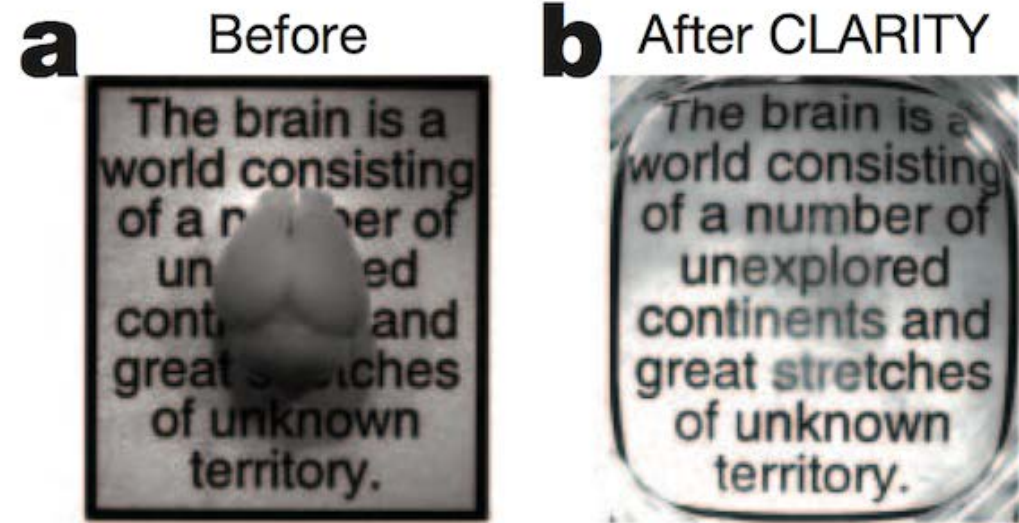
# Correcting scattering

Use NIR light, 2-photon microscopy



$$\sigma_s = \frac{2\pi^5 d^6}{3 \lambda^4} \left( \frac{n^2 - 1}{n^2 + 2} \right)^2$$

Remove lipids - Clarity



# Intensity calculations

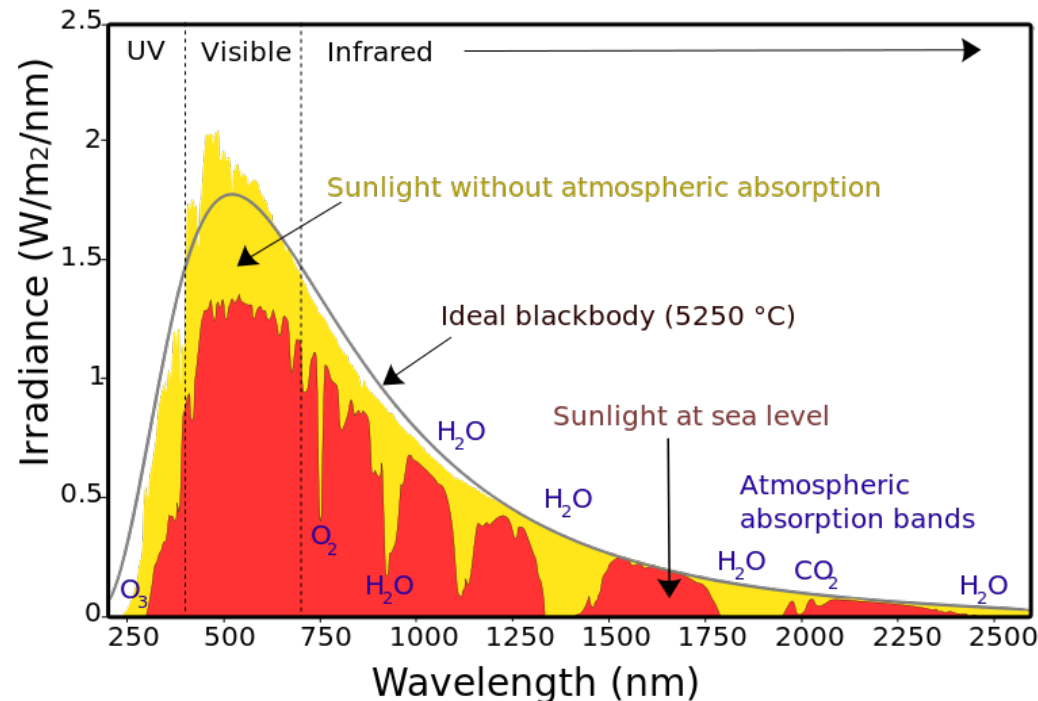
- Reporting light values used, it is imperative to use an intensity
- Energy/(s\*m<sup>2</sup>)

$I = \text{total power} / \text{total area}$

Shine 10 mW into 60x objective  
Field number = 22

What is intensity?  
How many photons?

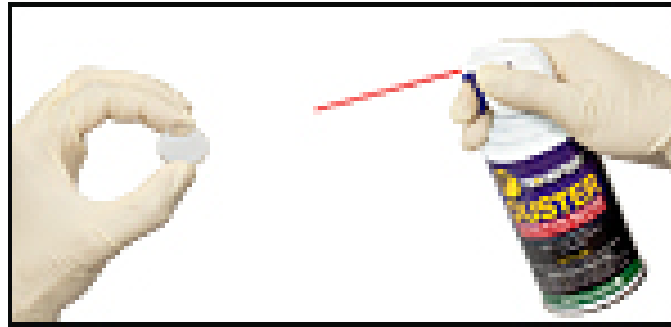
Spectrum of Solar Radiation (Earth)



Total solar intensity ~ 1000 W/m<sup>2</sup>  
On a bright day

~42% of that is visible light

# Cleaning optical components



Use lens or lint free paper



If you see dirt on your microscope

Try zooming up and down, is it before the sample, or after?

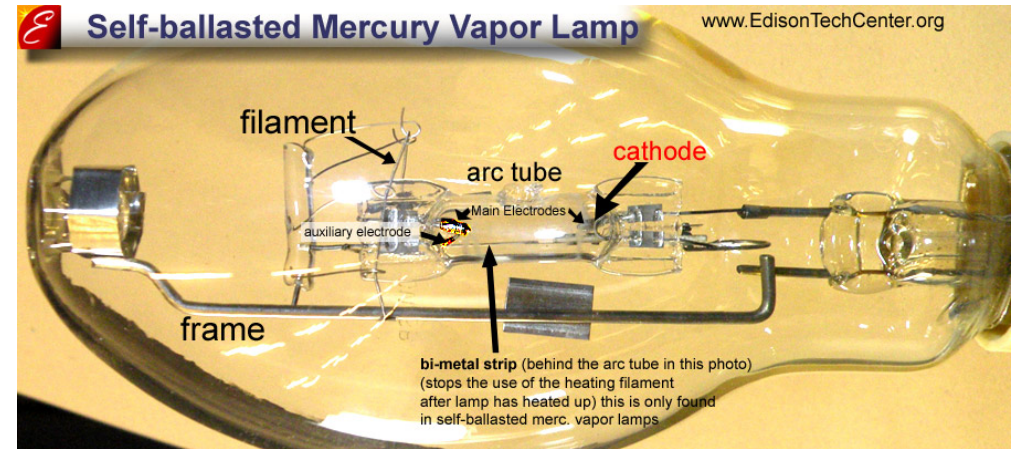
If it is imaged onto the camera, it must be somewhere close to an image plane



# Light sources

# Light source options

- Broadband sources (filaments)
- Light emitting diodes
- Lasers



# Properties of light

## Waveforms of Electromagnetic Radiation States

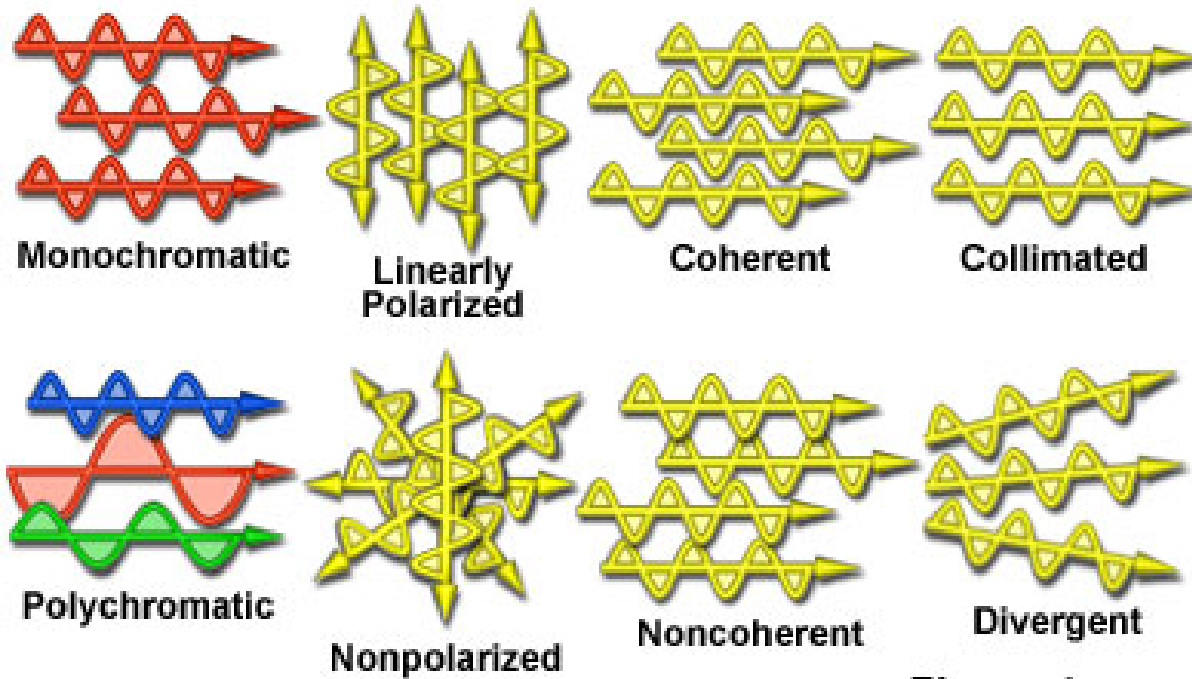


Figure 4

Other things to consider:

Intensity

Total cost

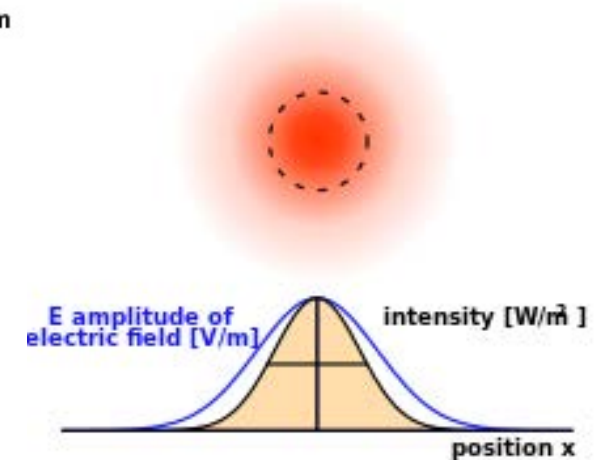
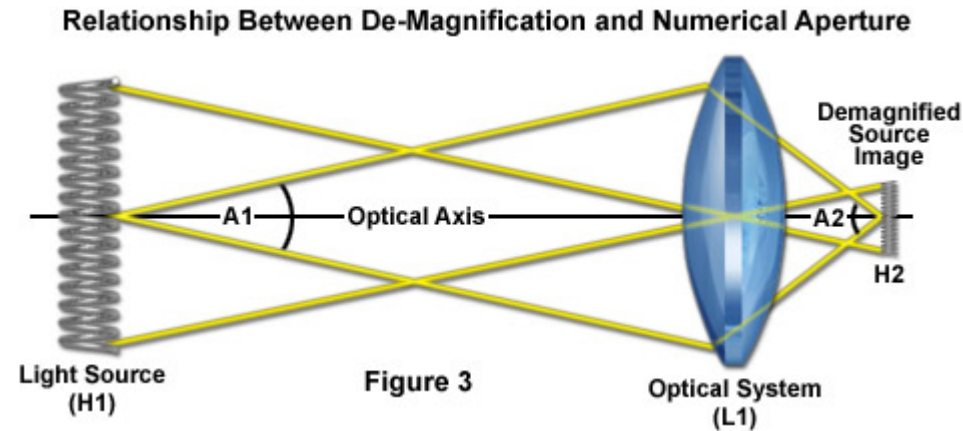
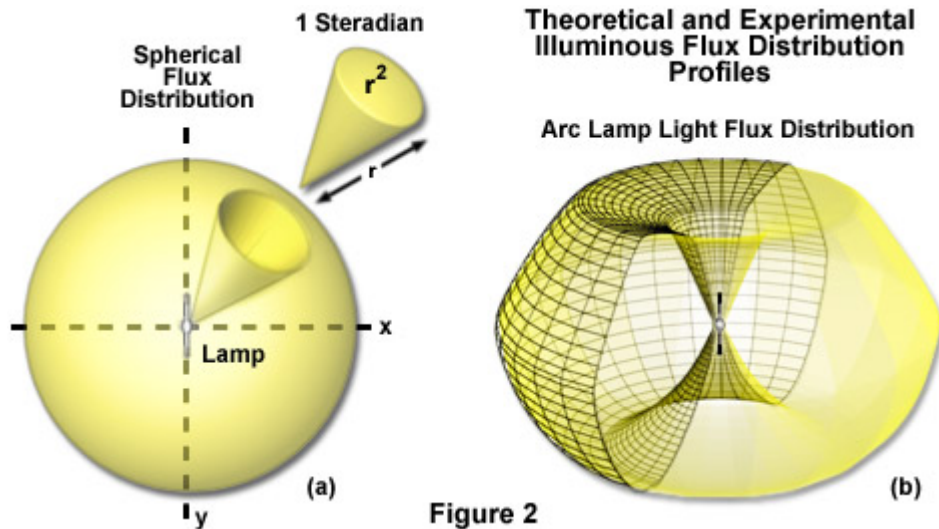
Compatibility with existing setup

Lifetime

Alignment

# The goal of illumination (mostly)

- Uniform, stable illumination across the entire field of view
- Lamps necessarily emit in all directions
- Even lasers emit Gaussian profile beams



# Importance of coherence

- Only coherent waves can interfere with each other
- Coherence has both a spatial and temporal scale
- Long spatial coherence will interfere of lenses, mirrors, specks of dust. We don't want that

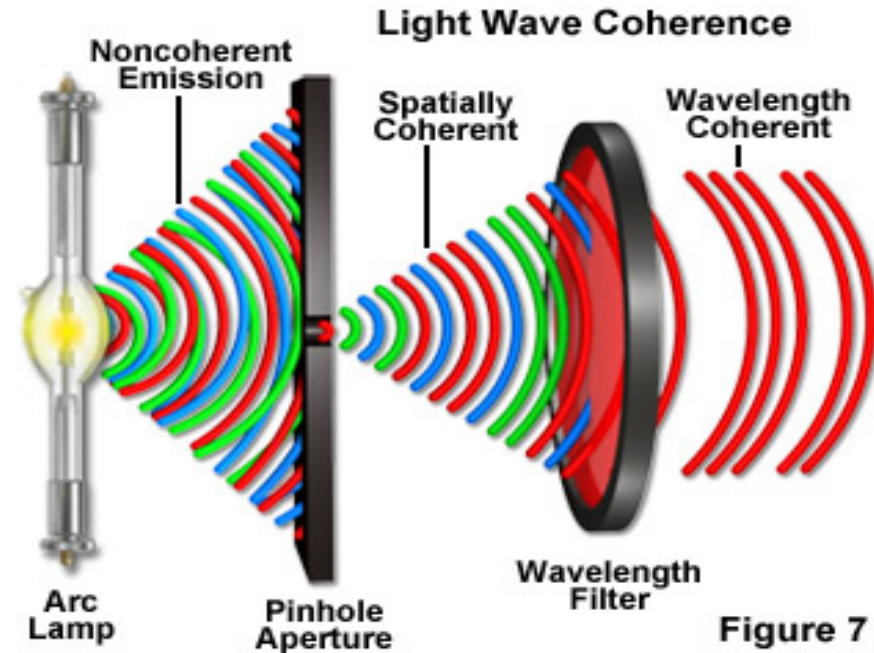
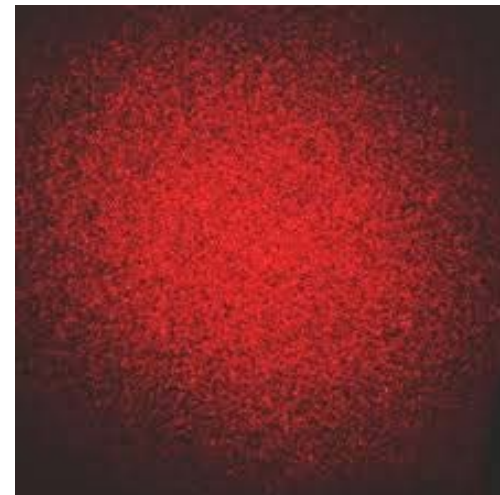


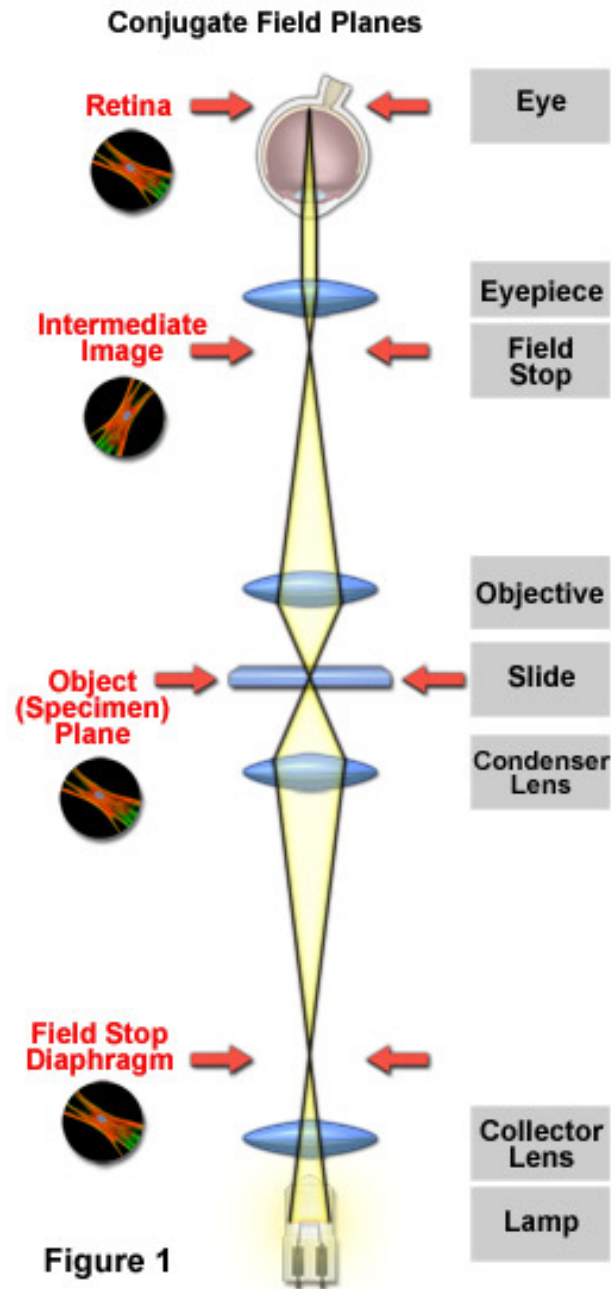
Figure 7



# Kohler illumination

- Ensure that sample is in infinity plane.
- Completely even illumination, don't have to worry about filament
- Form image of filament at focal plane of condenser
- As if it came from infinity

Conjugate Focal Planes in the Microscope for Köhler Illumination



On to Matlab...