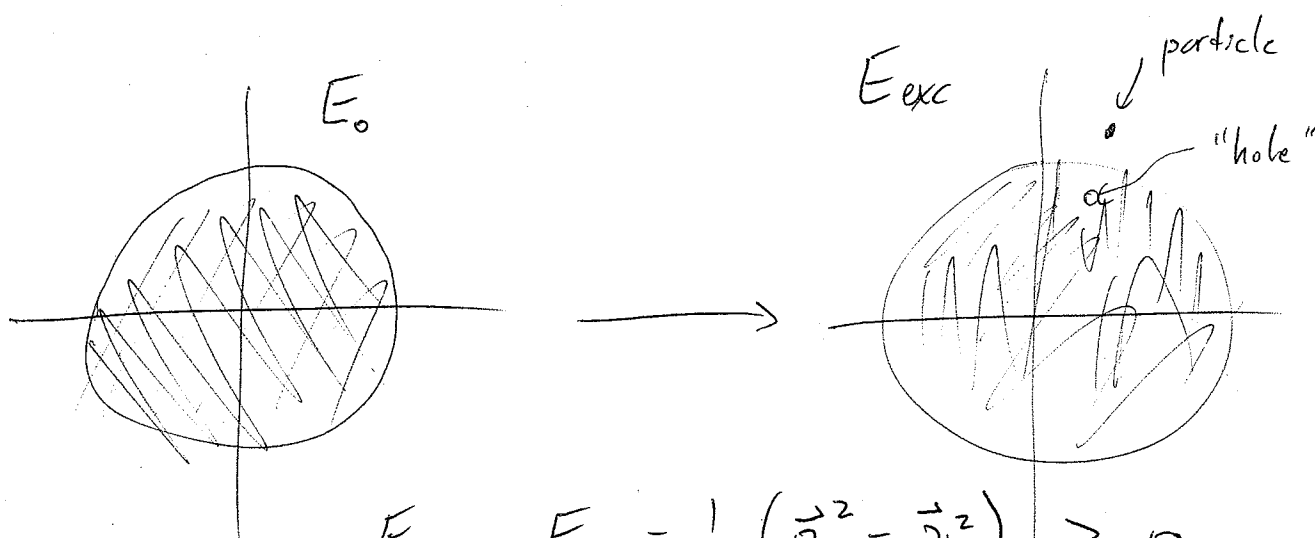


Free Fermi Gas - Nonzero Temperature

- For $T > 0$, excited states play a role. What are the low-energy excited states?

Particle-Hole Excitations



- $E_{exc} - E_0 = \frac{1}{2m} (\vec{p}_p^2 - \vec{p}_h^2) > 0$

- $E_{exc} - E_0 \rightarrow 0$ as $|\vec{p}_p|, |\vec{p}_h| \rightarrow p_F$

- $\Rightarrow$ Low-energy excitations are near the Fermi surface.

- Only particles & holes with $|E - E_F| \lesssim k_B T$ contribute significantly.

- For $T \ll T_F$, need to think QM'ically, in terms of Fermi gas. $T_F \sim 10^4$ K, so quantum mechanics important at room temp. & above!

Useful to work in grand canonical ensemble

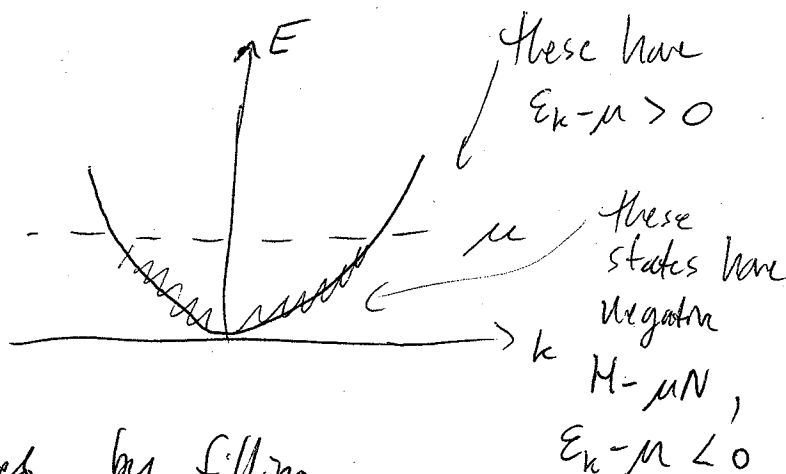
(let total # of electrons fluctuate $\leftrightarrow$ system can exchange electrons with a bath)

(call energies of this F , not E , since it's a kind of free energy)

- Then $H - \mu \hat{N}$ plays the role H played in canonical ensemble.

↑
number operator

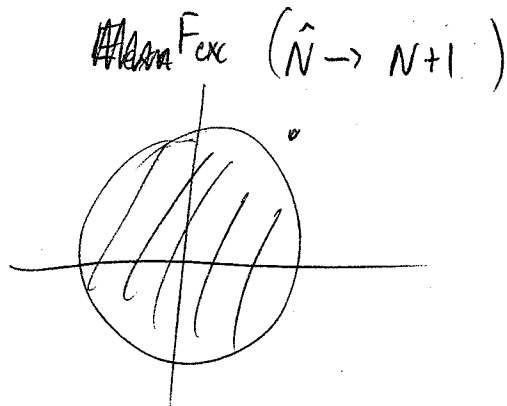
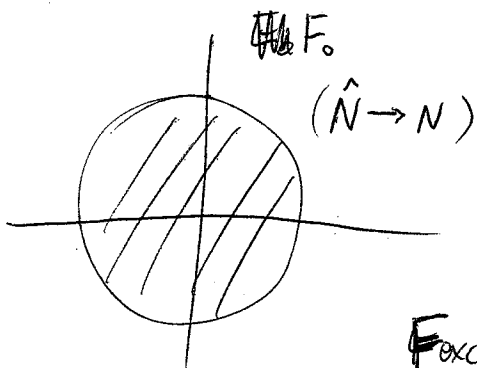
- To fill up fermi sea:



- Minimum $H - \mu N$ achieved by filling

states up to $E = \mu \Rightarrow \boxed{\mu(T=0) = E_F}$

- Single particle (or hole) excitations near fermi surface are low ^(free) energies:

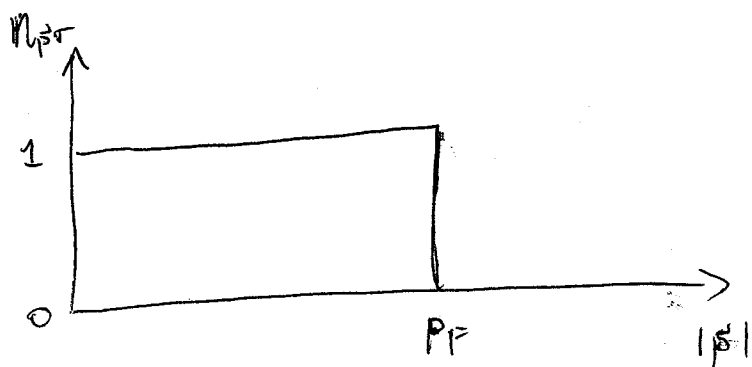


$$F_{exc} - F_0 = \frac{p_F^2}{2m} - \mu$$

Occupation Number — will be very useful at nonzero- T

$$n_{\vec{p}\sigma} = \Theta(p_F - |\vec{p}|) \text{ at zero-}T$$

- Probability state $|\vec{p}\sigma\rangle$ is occupied by an electron
 - Mean number of electrons in state $|\vec{p}\sigma\rangle$
- Exercise: Show these two quantities are the same.



• Total # of electrons: $N = \sum_{\vec{p}\sigma} n_{\vec{p}\sigma} \Rightarrow n = \sum_{\sigma} \int \frac{d^3\vec{p}}{(2\pi)^3} n_{\vec{p}\sigma}$

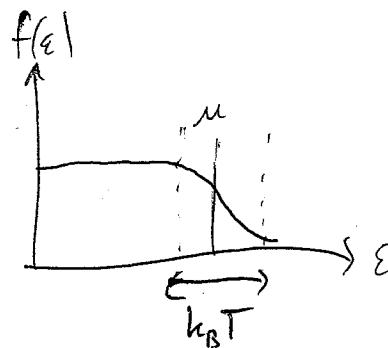
• Ground state energy: $E = \sum_{\vec{p}\sigma} \epsilon_{\vec{p}} n_{\vec{p}\sigma} \Rightarrow \mathcal{E} = \frac{E}{V} = \sum_{\sigma} \int \frac{d^3\vec{p}}{(2\pi)^3} \epsilon_{\vec{p}} n_{\vec{p}\sigma}$

Nonzero - T

- We will skip the derivation and write down the answer.
See any textbook on statistical mechanics (and also A&M ch. 2) for a derivation.

$$n_{p\sigma} = f(\epsilon_{\vec{p}}) = \frac{1}{e^{(\epsilon_{\vec{p}} - \mu)/k_B T} + 1}$$

↑
"Fermi function"



- Chemical potential is now a function of T :

Determined by:

$$N = 2 \sum_{\vec{k}} f(\epsilon_{\vec{k}}) \Rightarrow n = \int_{-\infty}^{\infty} d\epsilon D(\epsilon) f(\epsilon)$$

for spin

We need to know how to evaluate integrals like this

→ "Sommerfeld expansion"

- Won't go through whole derivation here. For details, see Appendix C of A&M.

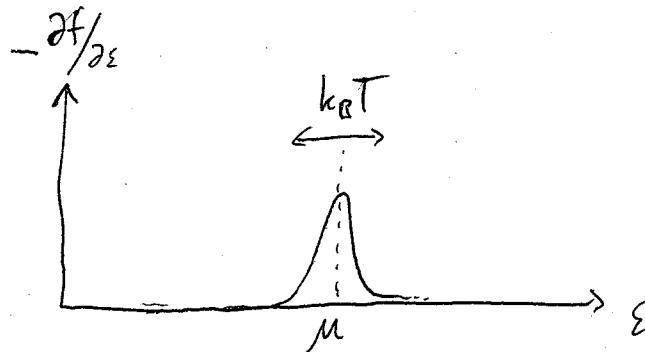
(5)

General problem: calculate $I = \int_{-\infty}^{\infty} d\varepsilon H(\varepsilon) f(\varepsilon)$

- $H(\varepsilon) \rightarrow 0$ for $\varepsilon \rightarrow -\infty$ (sufficiently fast)
- $H(\varepsilon) \lesssim C \varepsilon^m$ for large ε (some constant C , some integer m)

Want to exploit fact that $\frac{df}{d\varepsilon}$ is very small except

near $\varepsilon = \mu$:



Def. ne: $K(\varepsilon) = \int_{-\infty}^{\varepsilon} H(\varepsilon') d\varepsilon' \Rightarrow H(\varepsilon) = \frac{dK(\varepsilon)}{d\varepsilon}$

Integrate by parts: $I = \int_{-\infty}^{\infty} K(\varepsilon) \left(-\frac{\partial f}{\partial \varepsilon} \right) d\varepsilon$

Expand $K(\varepsilon)$: $K(\varepsilon) = K(\mu) + \sum_{n=1}^{\infty} \left[\frac{(\varepsilon - \mu)^n}{n!} \right] \left[\frac{d^n K(\varepsilon)}{d\varepsilon^n} \right]_{\varepsilon=\mu}$

- Crucial assumption: $K(\varepsilon)$ (and $H(\varepsilon)$) is analytic near $\varepsilon = \mu$.

Result:

$$\int_{-\infty}^{\infty} d\varepsilon H(\varepsilon) f(\varepsilon) = \int_{-\infty}^{\mu} H(\varepsilon) d\varepsilon + \frac{\pi^2}{6} (k_B T)^2 \left[\frac{dH}{d\varepsilon} \right]_{\varepsilon=\mu} + \mathcal{O}(T^4)$$

T-dependence of μ

$$N = \int_{-\infty}^{\mu} d\varepsilon D(\varepsilon) + \frac{\pi^2}{6} (k_B T)^2 \left[\frac{dD}{d\varepsilon} \right]_{\varepsilon=\mu} + \mathcal{O}(T^4)$$

Guess: $\mu(T) = \varepsilon_F + C(k_B T)^2 + (\text{higher order})$
 $\uparrow$ want to know this.

$$N = \underbrace{\int_{-\infty}^{\varepsilon_F} d\varepsilon D(\varepsilon)}_{=N} + \int_{\varepsilon_F}^{\varepsilon_F + C(k_B T)^2} d\varepsilon D(\varepsilon) + \frac{\pi^2}{6} (k_B T)^2 \left[\frac{dD}{d\varepsilon} \right]_{\varepsilon=\varepsilon_F}$$

$$\Rightarrow 0 = C(k_B T)^2 D(\varepsilon_F) + \frac{\pi^2}{6} (k_B T)^2 \left[\frac{dD}{d\varepsilon} \right]_{\varepsilon=\varepsilon_F}$$

$$\Rightarrow C = - \frac{\pi^2}{6 D(\varepsilon_F)} \left[\frac{dD}{d\varepsilon} \right]_{\varepsilon=\varepsilon_F} = \cancel{\frac{\pi^2}{12 \varepsilon_F}} - \frac{\pi^2}{12 \varepsilon_F}$$

$$\Rightarrow \boxed{\mu(T) = \varepsilon_F - \frac{\pi^2}{12} \frac{(k_B T)^2}{\varepsilon_F} + \dots}$$

small change
 $\sim \left(\frac{k_B T}{\varepsilon_F} \right) k_B T \ll k_B T.$

Specific Heat (a.k.a. Heat Capacity)

Defined by: $C_v = \frac{1}{V} \left(\frac{\partial E}{\partial T} \right)_{N,V} = \left(\frac{\partial u}{\partial T} \right)_{N,V}$
~~constant~~ energy density

- Significance: Measures how effective ΔT is in generating excitations of the system.
- Measurable: Put in heat $Q = \Delta E$, measure ΔT , $C_v = \frac{\Delta E}{\Delta T}$.

Let's calculate ~~u~~ $u(T)$:

$$u(T) = \int_{-\infty}^{\infty} d\varepsilon \varepsilon D(\varepsilon) f(\varepsilon) \quad (H(\varepsilon) = \varepsilon D(\varepsilon))$$

$$= \int_{-\infty}^{\mu} d\varepsilon \varepsilon D(\varepsilon) f(\varepsilon) + \frac{\pi^2}{6} (k_B T)^2 [\mu D'(\mu) + D(\mu)]$$

$$\approx \underbrace{\int_{-\infty}^{\varepsilon_F} d\varepsilon \varepsilon D(\varepsilon) f(\varepsilon)}_{u_0} + \int_{\varepsilon_F}^{\varepsilon_F + C(k_B T)^2} d\varepsilon \varepsilon D(\varepsilon) f(\varepsilon) + \frac{\pi^2}{6} (k_B T)^2 [\varepsilon_F D'(\varepsilon_F) + D(\varepsilon_F)]$$

$$= u_0 + [\varepsilon_F D(\varepsilon_F)] \left[-\frac{\pi^2}{6} (k_B T)^2 \frac{D'(\varepsilon_F)}{D(\varepsilon_F)} \right] + \frac{\pi^2}{6} (k_B T)^2 [\varepsilon_F D'(\varepsilon_F) + D(\varepsilon_F)]$$

$$= u_0 + \frac{\pi^2}{6} (k_B T)^2 D(\varepsilon_F)$$

$$\Rightarrow \boxed{C_v(T) = \frac{\pi^2}{3} k_B^2 T D(\varepsilon_F)}$$

again, D.O.S. at Fermi energy determines low-T physics!

• Comment: $C_v(T)$ much smaller than in classical gas, $C_v^{\text{class}}(T) = \frac{3}{2} k_B n \rightarrow \text{const. at low } T.$

Here, $C_v(T) = \frac{\pi^2}{2} \left(\frac{k_B T}{\epsilon_F} \right) n k_B$ ~~$\frac{\pi^2}{3} \left(\frac{k_B T}{\epsilon_F} \right) C_v^{\text{class}}(T)$~~

or $\frac{C_v(T)}{C_v^{\text{class}}(T)} \sim 10^{-1} - 10^{-2}$ at room temp!

→ Thermal excitations of electrons far from Fermi surface are frozen out.